



Separation of Delayed, Parameterized and Correlated Sources

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① Introduction

Galaxy Kinematics

Goal

② Proposed Method

Problem Formulation

Alternating Least Squares

Delay and Amplitude Estimation

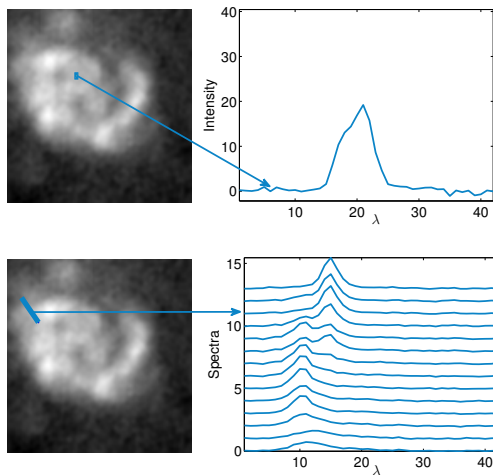
Numerical results

③ Conclusion



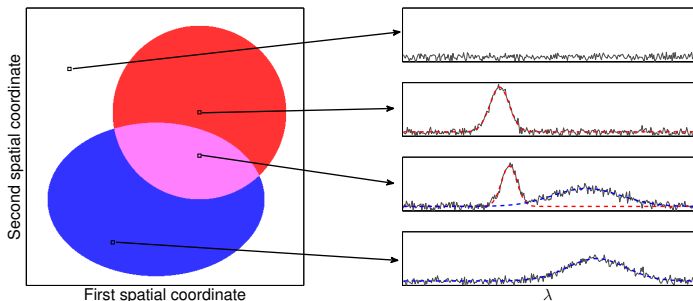
Spiral Galaxy M51-Nasa.gov

- Study of the internal gas motions
- Multiple structures



Galaxy NGC 4254

- Each structure is attributed to a peak
- Varying characteristics through the image



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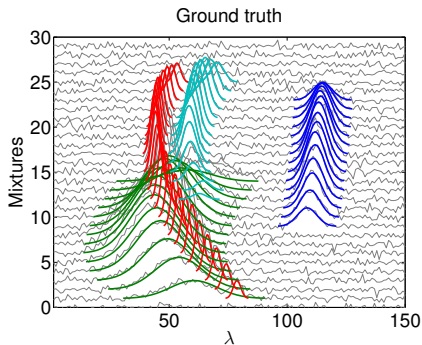
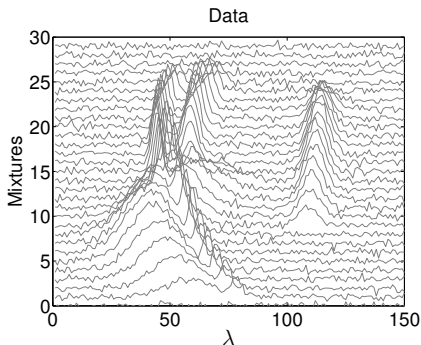
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3 Conclusion

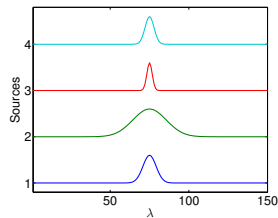
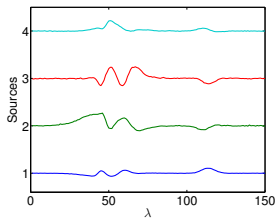
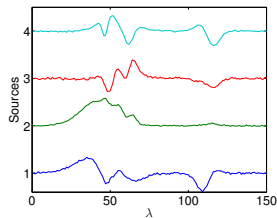
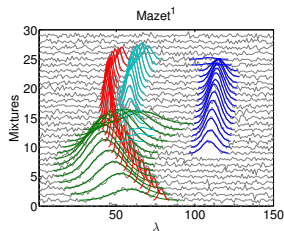
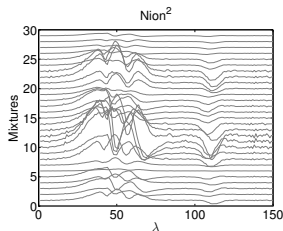
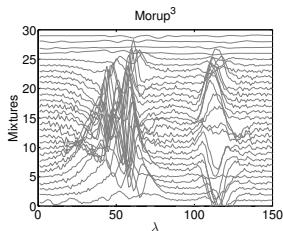
- Spectra i are assigned to the mixtures
- Peaks j are assigned to the sources
- Parameter estimation and peak tracking are done simultaneously



$I = 30$ mixtures $J = 4$ sources

Strong assumptions: independent, uncorrelated or orthogonal sources

Bayesian method



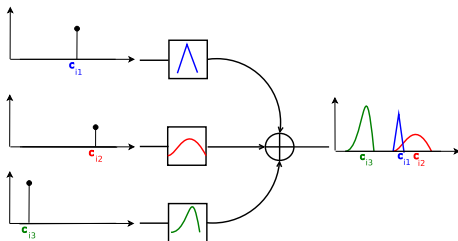
¹ Mazet et al., *Signal Processing*, 2015

² Nion et al., *LVA/ICA*, 2010

³ Morup et al., *ICA*, 2007

$$\mathbf{x}_i(\lambda) = \sum_{j=1}^J a_{ij} \mathbf{s}_j(\lambda - c_{ij}; w_j) + \mathbf{n}_i(\lambda) \quad \forall i$$

- Assumption: parameterized sources



- Special case: same waveform for all the sources

Estimate \mathbf{A} , \mathbf{C} and \mathbf{w} that minimize the residual error:

$$E(\mathbf{A}, \mathbf{C}, \mathbf{w}) = \sum_{i=1}^I \left\| \mathbf{x}_i(\boldsymbol{\lambda}) - \sum_{j=1}^J a_{ij} \mathbf{s}(\boldsymbol{\lambda} - c_{ij}; w_j) \right\|_2^2$$

$$\mathbf{A} = \begin{bmatrix} a_{11} & \dots & a_{1J} \\ \vdots & \ddots & \vdots \\ a_{I1} & \dots & a_{IJ} \end{bmatrix} \quad \mathbf{C} = \begin{bmatrix} c_{11} & \dots & c_{1J} \\ \vdots & \ddots & \vdots \\ c_{I1} & \dots & c_{IJ} \end{bmatrix} \quad \mathbf{w} = [w_1 \quad \dots \quad w_J]$$

ALS Scheme:

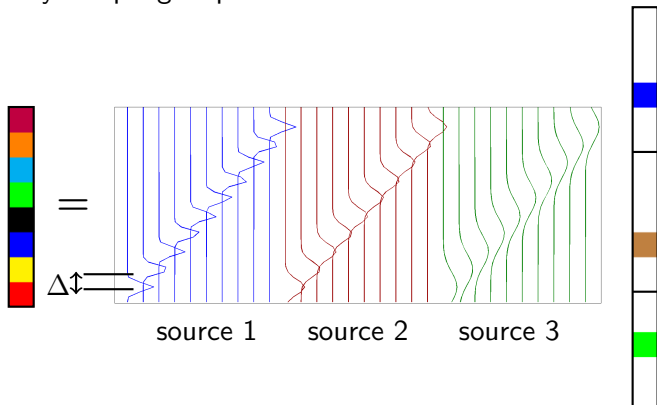
Until convergence:

- 1 minimize $E(\mathbf{A}, \mathbf{C}, \mathbf{w})$ w.r.t. \mathbf{w}
 - Levenberg-Marquardt algorithm
- 2 minimize $E(\mathbf{A}, \mathbf{C}, \mathbf{w})$ w.r.t. \mathbf{A} and \mathbf{C}
 - Parametric dictionary
 - OMP-like implementation

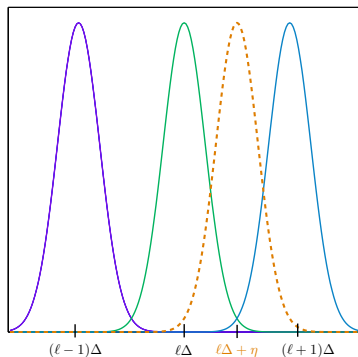
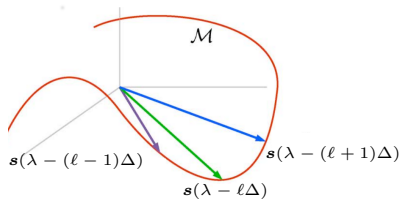
- Separable problem:

$$\min_{\mathbf{A}, \mathbf{C}} E(\mathbf{A}, \mathbf{C}, \mathbf{w}) \Leftrightarrow \min_{\mathbf{a}_i, \mathbf{c}_i} \left\| \mathbf{x}_i(\boldsymbol{\lambda}) - \sum_{j=1}^J a_{ij} \mathbf{s}(\boldsymbol{\lambda} - c_{ij}; w_j) \right\|_2^2 \quad \forall i$$

- Delay sampling step: Δ

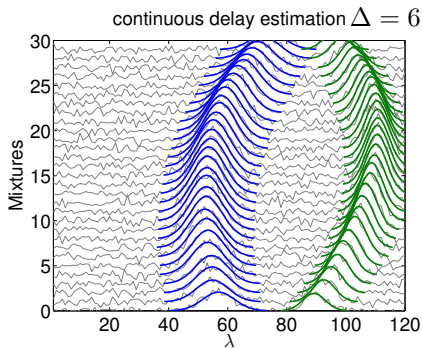
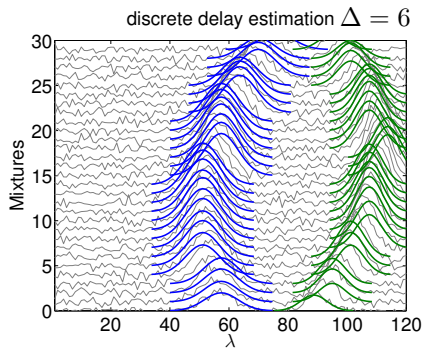


- Translation-invariant signals
- interpolation strategy: polar^{4,5}
- $c_{ij} = \ell\Delta + \eta \quad \ell \in \mathbb{Z}, |\eta| < \Delta/2$



⁴Ekanadham et al., *IEEE Trans. Signal Process.*, 2011

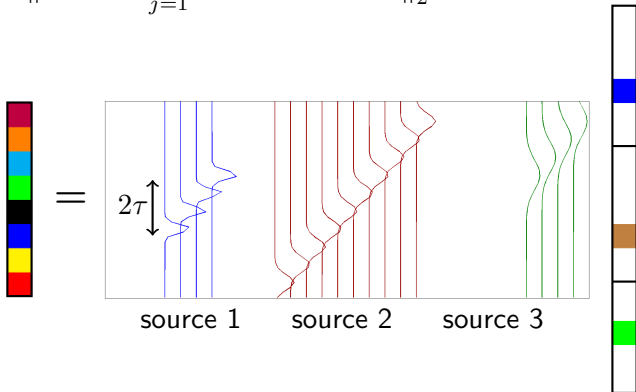
⁵Fyhn et al., *IEEE Trans. Signal Process.*, 2015

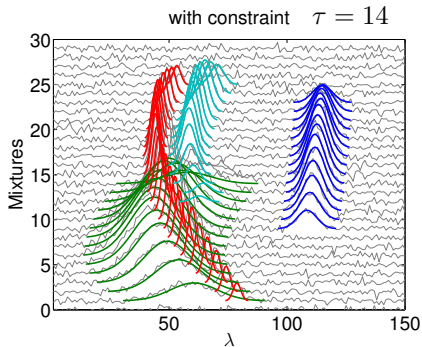
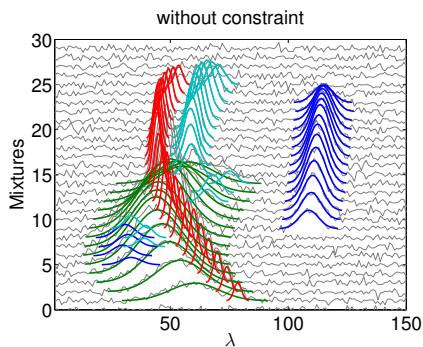


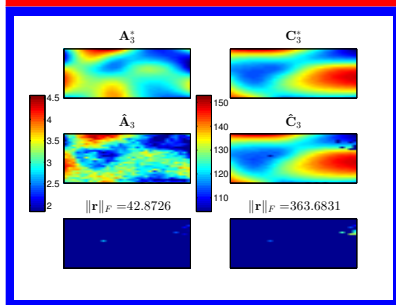
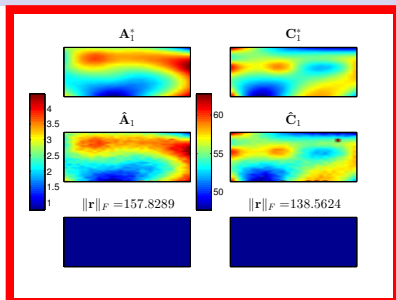
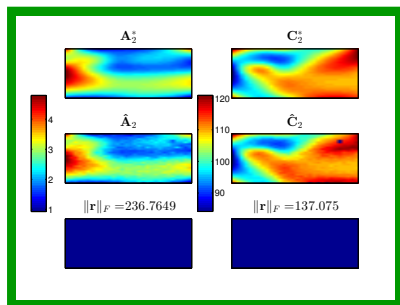
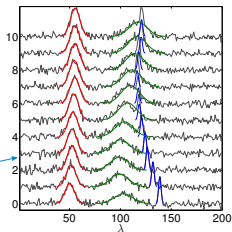
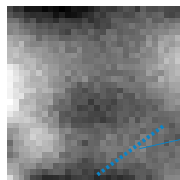
- Spatial neighboring information

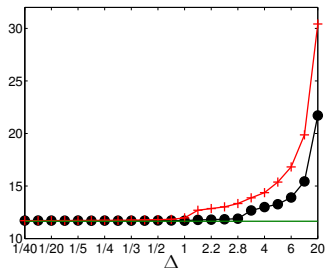
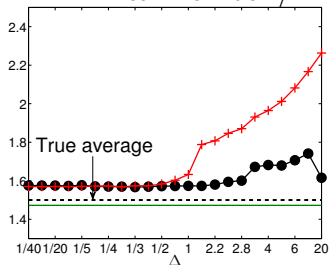
- $\bar{c}_{ij} = \frac{1}{\text{card}(\mathcal{V})} \sum_{v \in \mathcal{V}} c_{vj}$

- $\min_{\mathbf{a}_i, \mathbf{c}_i} \left\| \mathbf{x}_i(\boldsymbol{\lambda}) - \sum_{j=1}^J a_{ij} \mathbf{s}(\boldsymbol{\lambda} - c_{ij}; w_j) \right\|_2^2 \text{ s.t. } |c_{ij} - \bar{c}_{ij}| \leq \tau$

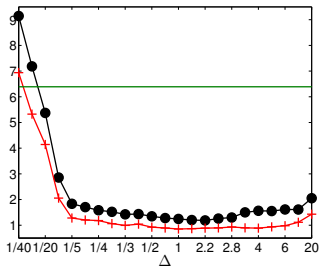






Residual norm E Peak number $/I$ 

Comp. time (seconds)



- $+$ Discrete version
- \bullet Continuous version
- $-$ MCMC

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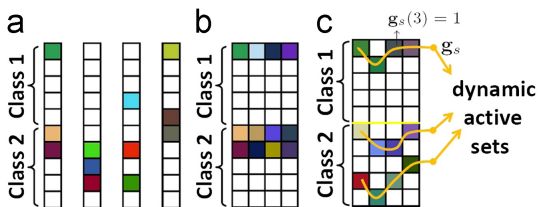
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3 Conclusion

- Delayed source separation model
- Parametrized and correlated sources
- Continuous delay estimation
- Constraint to ensure slow delay evolution
- As effective as the best competitors with much better computation time

- Varying source shape through the mixtures
- Dynamic joint sparse representation ⁶ + smooth regularization

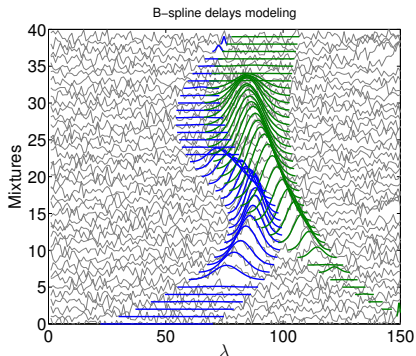


⁶Zhang et al., *Pattern Recognition*, 2012

- B-spline trajectories modeling

- $\hat{U} = \underset{U}{\operatorname{argmin}} \sum_{i=1}^I \left\| \mathbf{x}_i(\boldsymbol{\lambda}) - \sum_{j=1}^J \mathbf{s} \left(\boldsymbol{\lambda} - \sum_{k=1}^K u_{k,j} \mathbf{b}_{k,p}(i); w_{ij} \right) \right\|_2^2$

- $\hat{C} = B\hat{U}$



Thank you!

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